# Recent Results from the PandaX-II Experiment

Andi Tan (University of Maryland) on behalf of the PandaX-II Collaboration Osaka University, Dec. 21, 2017

International workshop on "Axion physics and dark matter cosmology"



- Dark Matter Direct Detection
- PandaX at CJPL
- Recent Results from the PandaX-II
- Future of PandaX
- Conclusion









# Dark Matter Detection

 $\mathbf{\mathcal{O}}$ 

#### Indirect Search



AMS

#### DAMPE

#### **Direct Search**

SM

DM

#### Collider Search

**Overall view of the LHC experiments.** 





SM

DM







15 JUNE 1985

1 JULY 1985

 $= m_2/m_1$ 

PHYSICAL REVIEW D

VOLUME 55, NUMBER 1

#### **VOLUME 31, NUMBER 12**

Detectability of certain dark-matter candidates

Mark W. Goodman and Edward Witten Joseph Henry Laboratories, Princeton University, Princeton, New Jersey 08544 (Received 7 January 1985)

#### PHYSICAL REVIEW LETTERS

#### **Bolometric Detection of Neutrinos**

Blas Cabrera, Lawrence M. Krauss, and Frank Wilczek Department of Physics, Stanford University, Stanford, California 94305 Lyman Laboratory of Physics, Harvard University, Cambridge, Massachusetts 01238 Institute for Theoretical Physics, University of California, Santa Barbara, California 93106 (Received 14 December 1984)

$$E_{1} = \frac{E_{0}}{(1+A)^{2}} (\cos\theta \pm \sqrt{A^{2} - \sin^{2}\theta})^{2}$$

$$T = E_2 = \frac{4A}{(1+A)^2} E_0 \cos^2 \phi$$
 A

IMPs and Neutrons scatter from the scatter from the Atomic Electrons

### Direct Detection

- DM: velocity  $\sim 1/1500$  c, mass  $\sim 100$  GeV, KE ~ 20 keV
- Nuclear recoil (NR): recoiling energy ~10 keV
- Electron recoil (ER): 10-4 suppression in energy, very difficult to detect
- New ideas exist, e.g. Hochberg, Zhao, and Zurek, PRL 116, 011301











### Available Space and Global Racing



Nature Physics 13, 212–216 (2017)



- Event rate approaches a few cnts/100kg/y
- Lowest exclusion at ~8×10<sup>-47</sup>cm<sup>2</sup> at ~40GeV/c<sup>2</sup>
- ~2 orders away from the neutrino floor







# invisible phonons ecoil

#### Dual Phase Xenon Experiment



• Dual phase TPC technology allows:

- 1. 3D positioning
- 2. Single Scatter selection

S1 and S2 ratio allows ER/NR discrimination











#### Need three legs to make a table stable





250 kg, concluded 2016, Sanford Lab LZ(7-ton) in preparation

Largest LXe TPC ever built cylinder: 96 × 97cm active Xe target 2.0t(3.2t total) 248 PMTs (Hamamatsu R11410-21)

- World leading bkg level: 0.2×10<sup>-3</sup> evt/day/kg/keV
- First SR: 1024 kg x 34.2 day, no candidate found
- Minimum limit: 7.7 x 10<sup>-47</sup> cm<sup>2</sup> @ 35 GeV











### China Jin-Ping underground Lab







#### XMASS

Usable xenon: 835 kilograms Status: Reported 6.7 days of data. Plans for a 1.5-tonne experiment in 2014 at a cost of US\$12 million.

#### XENON100

Usable xenon: 62 kilograms Status: Reported 225 days of data. Construction begins in 2013 for \$12-million tonne-scale experiment.

#### LUX

Usable xenon: 350 kilograms Status: Taken surface data and has just started below ground. Plans for multi-tonne experiment in 2016-17, at a cost of \$30 million.

#### PANDAX

Usable xenon: 120 kilograms Status: Yet to take data. Plans for tonne-scale experiment in 2016 at a cost of \$15 million.

XMASS: Xe detector for weakly interacting massive particles; LUX: Large Underground Xenon detector: PANDAX: Particle and Astrophysical Xenon Time Projection Chamber

### China Jin-Ping underground Lab

#### Dark-matter hunt gets deep

China launches world's deepest particle-physics experiment — but it joins a crowded field.

#### **Eugenie Samuel Reich**

20 February 2013 | Corrected: 21 February 2013



Ongoing experiments in Italy, the United States and Japan are now being joined by a fourth in China, called PandaX (see 'Dark and deep'). Installed in the deepest laboratory in the world, 2,500 metres under the marble mountain of JinPing in Sichuan province, PandaX will this year begin monitoring 120 kilograms of xenon. The team hopes to scale the tank up to 1 tonne by 2016, which would mean that the experiment had developed more quickly than any other dark-matter search. "We want to demonstrate that world-class research in dark matter is possible in China," says Xiangdong Ji, a physicist at Shanghai Jiao Tong University in China and a spokesman for PandaX.

2,500 m

#### nature International weekly journal of science

#### NATURE | NEWS











# PandaX Collaboration

#### Started in 2009, ~50 people



- Shanghai Jiao Tong University (2009-)
- Peking University (2009-)
- Shandong University (2009-)
- Shanghai Institute of Applied Physics, CAS (2009-)
- University of Science & Technology of China (2015-)
- China Institute of Atomic Energy (2015)
- Sun Yat-Sen University (2015-)
- Yalong Hydropower Company (2009-)
- University of Maryland (2009-)
- Alternative Energies & Atomic Energy Commission (2015-)
- University of Zaragoza (2015-)
- Suranaree University of
  - Technology(2016-)







# PandaX Experiments





PandaX-I: 120 kg DM experiment

PandaX-II: 580 kg PandaX-xT: DM experiment multi-ton DM experiment







PandaX-III: 200 kg to 1 ton HP gas <sup>136</sup>Xe **OvDBD** experiment

**PANDAX** = Particle and Astrophysical Xenon Experiments







# The PandaX-II Detector

- New inner vessel with clean SS
- New and taller TPC
- More 3" PMTs and improved base design with split ±HV power supply
- New isolated skin veto region







Nov. 2016 – Mar. 2017, 2<sup>nd</sup> distillation campaign and recommissioning 2016 2017 Apr.22 – July15, dark Jul – Oct, ER calibration matter data taking & tritium removal (Run10, 77.1 days)

2015

Mar. 9 – June 30, low background with 10-fold reduction of Kr (Run9, 79.6 days) Nov. 22 – Dec. 14, Physics commission (Run8, 19.1 days, stopped due to high Krypton background)

- Run9 = 79.6 days, exposure: 26.2 ton-day
- Largest reported DM exposure to date

# PandaX-II Run History

• Run10 = 77.1 days, exposure: 27.9 ton-day





### Run8+9(33 t-day) SI and SD Results





#### Minimum elastic SI exclusion: 2.5x10-46 cm<sup>2</sup> @ 40 GeV/c<sup>2</sup> PRL 117, 121303 (2016)



Minimum  $\chi$ -n SD cross section limit: 4.1x10<sup>-41</sup> cm<sup>2</sup> at 40 GeV/c<sup>2</sup> PRL 118, 071301 (2017) 🔄 PANDAX

# PandaX-II Run9 Axion and Inelastic Results

PRL 119, 181806 (2017)



Among the leading axion search on axionelectron coupling using DD experiments





- Opened up energy window to access initial-final mass difference up to 300 keV (high mass DM, ~TeV)
- Tightest direct constraint on this to date (to be published) PRD 96, 102007 (2017)



### 2017 new data and results







### New SI DM search results from Run10

- Improved trigger threshold (from 80 to 50 PE $\approx$ 2 electrons)
- Calibration
- Channel-by-channel SPE efficiency ( $\mathcal{E}_{ZLE}$ ) Improved detector ER/NR response model • 2.5 times reduction in total **background** 
  - Kr85  $\downarrow$ 6 times
  - Accidental J 3 times
  - Xe127 ↓13 times







# Calibration Program

- Gain Calibration
- ZLE Efficiency on S1 and S2
- NR Calibration using <sup>241</sup>AmBe
- ER calibration using CH<sub>3</sub>T

Uniformity correction and Energy reconstruction







#### Cali: Overall ZLE efficiency by LED Vs. S1



- Average efficiency measured at 3 PE (S1) threshold) was about 80% using blue LED
- PMT Double-PE emission would further improve the efficiency
- Little impact to S2, little impact to position  $\bullet$ reconstruction











### Cali: Uniformity Correction

- Z correction on S2
  - electron absorption to electronegative impurities
- X-Y correction on S2
  - geometry
  - Field non-uniformity
- 3D correction on S1
  - geometry
  - Rayleigh Scatter
  - Light attenuation







# Cali: Election Lifetime



- $\bullet$
- Significantly improved from Run9

Electron lifetime on average 800µs (1.4 m drift distance) in Run10, and generally stable











# 

#### Cali: Energy Reconstruction



$$E_{ee} = W \times \begin{pmatrix} S1 & S2 \\ PDE \end{pmatrix} + EEE \times SE$$

For this analysis

 SEG determined with ZLE efficiency taken into account Utilized a more careful treatment for the S2 saturation • Resulting best fit <2% with expected energies (· · ·







# Cali: ER and NR



Neutron calibration: AmBe source deployed



 ER calibration using tritiated methane (pioneered by LUX)

 Selected data with electron lifetime ~700 s, ~8000 low energy ER events







### Data/model comparison(example)



- A tuning of the N<sub>ex</sub>/N<sub>i</sub> (excitation/ionization) parameter was made on the NEST model, after which data and MC yield good agreement
- Consistent efficiencies obtained using ER/NR calibrations









#### Cali: Direct comparison of ER and NR bands



• Events leaked below the NR median: 0.53(8)%

![](_page_24_Picture_3.jpeg)

![](_page_24_Picture_5.jpeg)

![](_page_24_Picture_9.jpeg)

# Calibration to Background

![](_page_25_Figure_1.jpeg)

Game of Thrones: Background level We want keep as clean as possible in this region and waiting for DM scattering events

![](_page_25_Picture_3.jpeg)

40

![](_page_25_Picture_5.jpeg)

![](_page_25_Picture_6.jpeg)

# BKG: 2<sup>nd</sup> distillation campaign

- After ER calibration, realized that the getter could not remove tritium background effectively
- Suspected tritium attached to wall, emanation rate balance with removal rate
- 2<sup>nd</sup> distillation campaign (for Kr and tritium)
- Nov. Mar 2017: recuperate distillation refill, flush (closed) detector with warm xenon

#### First beneficial occupancy of CJPL-II!

![](_page_26_Picture_6.jpeg)

![](_page_26_Picture_7.jpeg)

27

![](_page_26_Picture_9.jpeg)

![](_page_26_Picture_10.jpeg)

# Background level

	Run9	Run10	
Xe127	0.42	0.021	
Tritium	0	0.27	
Kr85	1.19	0.20	
Rn222	0.13	0.12	
Rn220	0.01	0.02	
Detector ER	0.20	0.20	
Solar neutrino	0.01	0.01	
Xe136	0.0022	0.0022	
Total	1.96	0.79	

Original <sup>127</sup>Xe gone, additional introduced by a bottle of surface xenon during distillation

Based on best fit to data (later)

Rest are consistent between Run9 and Run10

0.8 mDRU ~ 2 events a day!

![](_page_27_Picture_6.jpeg)

![](_page_27_Picture_8.jpeg)

![](_page_27_Picture_12.jpeg)

### BKG: Energy spectrum in Run 10

![](_page_28_Figure_1.jpeg)

Data and expected background in good agreement

![](_page_28_Figure_3.jpeg)

![](_page_28_Picture_5.jpeg)

![](_page_28_Picture_7.jpeg)

![](_page_28_Picture_11.jpeg)

### BKG: Distribution of events (run10)

![](_page_29_Figure_1.jpeg)

	ER	Accident al	Neutron	Total Fitted	To Obs
Run9	376.1	13.5	0.85	390±50	3
Below NR Median	2.0	0.9	0.35	3.2±0.9	
Run10	172.2	3.9	0.83	177±33	1
Below NR Median	0.9	0.6	0.33	<b>1.8±0.5</b>	

Appears to have a downward fluctuation of background!

![](_page_29_Picture_4.jpeg)

![](_page_29_Picture_6.jpeg)

![](_page_29_Picture_7.jpeg)

### WIMP-nucleon SI result

![](_page_30_Figure_1.jpeg)

- Lowest exclusion at 8.6×10-47 cm<sup>2</sup> at 40GeV/c<sup>2</sup>

![](_page_30_Figure_5.jpeg)

 most stringent limit for WIMP-nucleon cross section for mass >100GeV Improved from PandaX-II 2016 limit about 2.5 time for mass>30 GeV

![](_page_30_Picture_7.jpeg)

![](_page_30_Picture_9.jpeg)

![](_page_30_Picture_13.jpeg)

### PandaX Future

![](_page_31_Picture_1.jpeg)

![](_page_31_Picture_3.jpeg)

![](_page_31_Picture_7.jpeg)

### PandaX new home: CJPL-II

![](_page_32_Figure_1.jpeg)

![](_page_32_Figure_2.jpeg)

• Dark matter, 0vDBD, nuclear astrophysics, low background CFT

**B2, PandaX site!** 

![](_page_32_Picture_5.jpeg)

![](_page_32_Picture_7.jpeg)

![](_page_32_Picture_11.jpeg)

![](_page_32_Picture_12.jpeg)

![](_page_33_Picture_1.jpeg)

- Preparing new experiments in CJPL-II, hall #B2
- Intermediate stage:
- PandaX-4T (4-ton target) with SI sensitivity  $\sim 10^{-47} \text{ cm}^2$
- On-site assembly and commissioning: 2019-2020
- Eventual goal: G3 xenon dark matter detector (~30T) in CJPL to "neutrino floor" sensitivity

### PandaX-xT Experiment

![](_page_33_Picture_8.jpeg)

![](_page_33_Picture_10.jpeg)

### PandaX-III: High pressure <sup>136</sup>Xe TPC

- ODBD signal: two electrons emitting from the same vertex with a summed energy at the Q value (tracking essential)
- TPC: 200 kg, 10 atm, symmetric, double-ended charge readout plane with micromegas module with cathode in the middle
- Four more upgraded modules for a ton scale experiment
- Published CDR recently: <u>ArXiv:1610.088883</u>

![](_page_34_Figure_5.jpeg)

![](_page_34_Figure_6.jpeg)

![](_page_34_Figure_7.jpeg)

![](_page_34_Picture_8.jpeg)

### Conclusion

#### Searching for WIMPs is far from over

- PandaX experiment since 2012 has been gone through two generations of detectors
- PandaX-II has reached and remains at the forefront of the DM search, and will continue data taking till end of 2018
- The collaboration is going forward in preparation for PandaX-III and PandaX-xT.

![](_page_35_Picture_6.jpeg)

![](_page_35_Picture_8.jpeg)

![](_page_35_Picture_12.jpeg)